

Relativistic Corrections for Time and Frequency Transfer in Optical Fibres

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Relativistic geodesy: first steps towards a new geodetic technique
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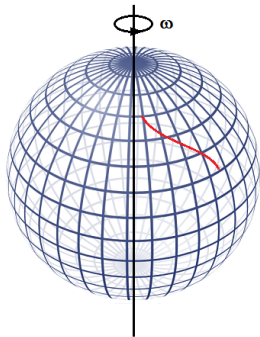


Project background and motivation

- EMRP project International Timescales with Optical Clocks (ITOC)
- Comparisons of distant optical clocks
- Theoretical description of TT and FT with accuracy required by optical clocks

Objectives

- Fully relativistic description of signal propagation in optical fibres
- Deriving formulas for signal propagation time in optical fibre with 1 ps accuracy
- Deriving formulas for frequency change during signal propagation in optical fibre with relative accuracy of 10^{-18}

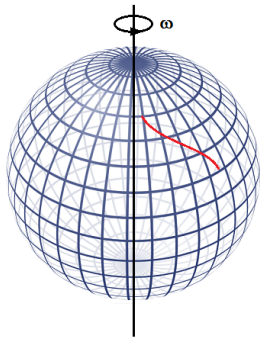


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Outline

- 1 Fibre time transfer
- 2 Fibre frequency transfer

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Fibre time transfer [Gersl et al., 2015]

$$T_{AB} = \frac{1}{c} \int_0^L n \, dl + \frac{1}{c^2} \int_0^L \mathbf{v} \cdot \mathbf{s}_l \, dl + \frac{1}{c^3} \int_0^L n (W + v^2/2) \, dl + O(c^{-4})$$

where n is the fibre effective refractive index, \mathbf{v} is the velocity vector field of the fibre in the GCRS frame, \mathbf{s}_l is the tangent vector field of the fibre, $v^2 = \mathbf{v} \cdot \mathbf{v}$, and all quantities are evaluated at $t = t_B$.

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Fibre time transfer: the Sagnac term

- for two-way time transfer only the Sagnac term remains:

$$\frac{1}{2}(T_{34} - T_{12}) = \frac{1}{c^2} \int_0^L \mathbf{v} \cdot \mathbf{s}_l \, dl$$

- Let's introduce a **Terrestrial frame** which rotates rigidly with respect to the spatial part of GCRS with angular velocity vector $\boldsymbol{\omega}$ such that:

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$$\frac{1}{2}(T_{34} - T_{12}) = \frac{2\omega A}{c^2} + \frac{1}{c^2} \int_0^L \mathbf{v}_R \cdot \mathbf{s}_l \, dl$$

where A is the area of the projection of the fibre on the equatorial plane, when connecting both end with the rotation axis.

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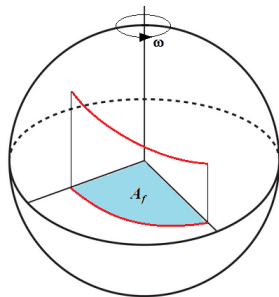
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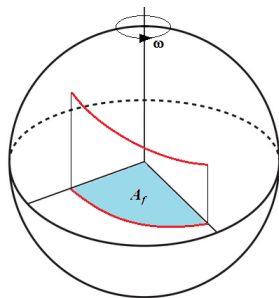
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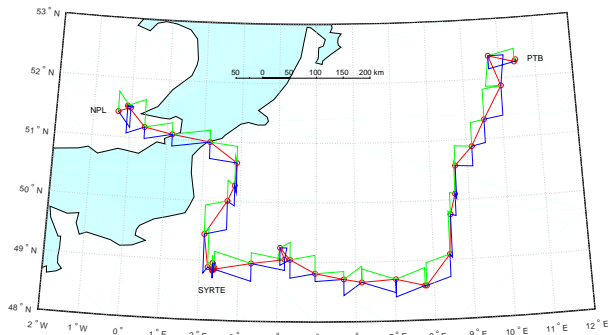
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Sagnac delays in the REFIMEVE+ network



fibre link	Length/km	Correction/ps
PTB-SYRTE	1401	3976 ± 27
NPL-SYRTE	813	1214 ± 6

Parameters uncertainty for fibre time transfer

Parameter	Uncertainty
Fibre length (1-way only)	0.2 mm
Refractive index (1-way only)	3×10^{-10}
Fibre endpoints position	200 m
Fibre inner points position	600 m
Fibre velocity in co-rotating frame	9 cm/s
Earth angular velocity	~ 0.01 % (relative)
Gravitational plus centrifugal potential (1-way only)	~ 30 % (relative)

Table : Input parameters and their maximal uncertainties sufficient for 1 ps uncertainty in time transfer. The values were obtained for situations where the sensitivity of a correction to a parameter is maximized and they are calculated for 1000 km long fibre [Gersl et al., 2015].

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Clock frequency comparison

- Determination of the frequency ratio f_A/f_B between the proper frequencies f_A and f_B delivered by clocks A and B
- In practice this is achieved using a transmission of photons from A to B and the formula

$$\frac{f_A}{f_B} = \frac{f_A}{\nu_A} \frac{\nu_A}{\nu_B} \frac{\nu_B}{f_B},$$

where ν_A = proper frequency of the photon as measured on A at instant of emission t_A , and ν_B = proper frequency of the same photon on B at t_B (instant of reception).

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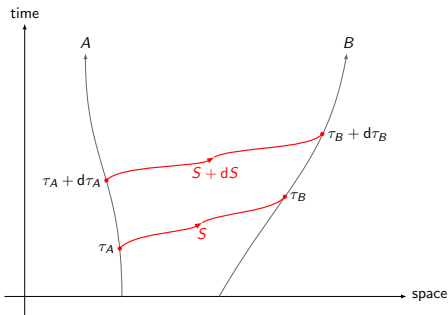
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Fundamental relations of frequency transfer



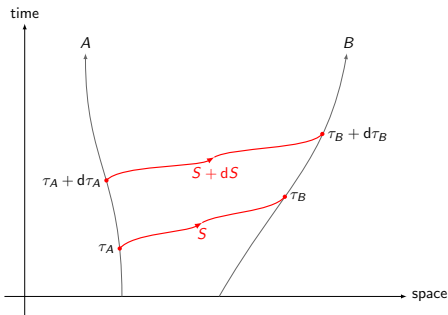
- The proper frequency measured by $M = \{A, B\}$ is:

$$\nu_M = \frac{1}{2\pi} \frac{dS}{d\tau_M} \quad (1)$$

- Introducing the covariant wave vector $k_\alpha^M = (\partial_\alpha S)_M$ and the four-velocity $u_M^\alpha = dx_M^\alpha / d\tau_M$:

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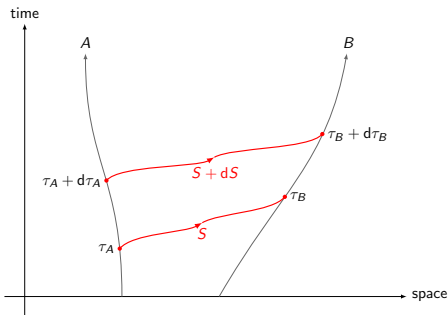
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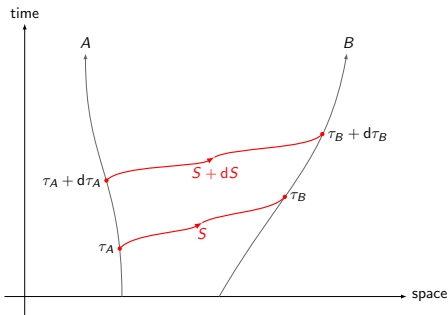
- Introducing $v^i = dx^i/dt$ and $\hat{k}_i = k_i/k_0$, it is usually written as:

$$\frac{\nu_A}{\nu_B} = \frac{u_A^0}{u_B^0} \frac{k_0^A}{k_0^B} \frac{1 + \frac{\hat{k}_i^A v_A^i}{c}}{1 + \frac{\hat{k}_i^B v_B^i}{c}}$$

- from (1) we deduce:

$$\begin{aligned} \frac{\nu_A}{\nu_B} &= \frac{d\tau_B}{d\tau_A} \\ &= \left(\frac{d\tau}{dt} \right)_A^{-1} \left(\frac{d\tau}{dt} \right)_B \frac{dt_B}{dt_A} \end{aligned}$$

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Splitting of the terms

- One term depends only on the state (velocity and location) of the receiver and emitter

$$\frac{u_A^0}{u_B^0} = \left(\frac{d\tau}{dt} \right)_A^{-1} \left(\frac{d\tau}{dt} \right)_B$$

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The proper time / coordinate time relation in GCRS

- The scalar potential $W(X^\alpha) \equiv W(cTCG, \mathbf{X})$ of the GCRS metric can be split into three parts: contribution of the Earth itself, of tidal forces, and of inertial forces [Soffel et al., 2003]

$$W = W_E + W_{\text{tidal}} + W_{\text{inertial}}$$

- Then the relation between proper time and TCG is [Wolf and Petit, 1995]:

$$\frac{d\tau}{dTCG} = 1 - \frac{1}{c^2} \left[W_E(X^\alpha) + \frac{v^2}{c^2} + \bar{W}(x_E^\alpha + X^\alpha) - \bar{W}(x_E^\alpha) - \bar{W}_{,k}(x_E^\alpha) X^k \right]$$

where \bar{W} is the Newtonian potential of external masses, and $(x^\alpha) \equiv (cTCB, \mathbf{x}^i)$ are the BCRS coordinates.

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One-way fibre propagation term

$$\frac{dt_B}{dt_A} = 1 + \frac{1}{c} \int_0^L \left(\frac{\partial n}{\partial t} + n\alpha \frac{\partial T}{\partial t} \right) dl + \frac{1}{c^2} \int_0^L \frac{\partial \mathbf{v} \cdot \mathbf{s}_l}{\partial t} dl$$

where T is the temperature and other quantities are defined in the section on time transfer [Gersl et al., 2015]

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Fibre propagation

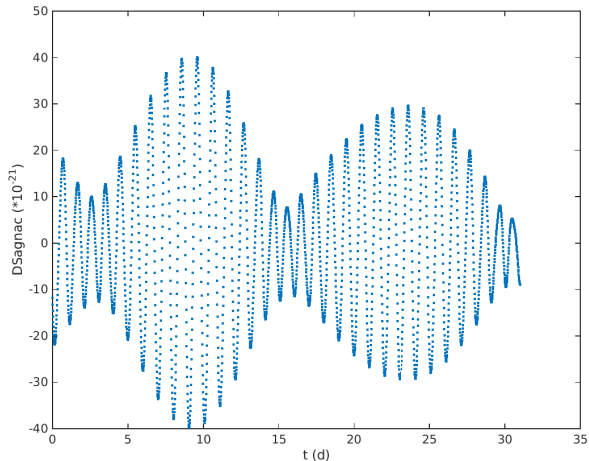
Magnitudes of the different terms

Effect	Correction
Variations in length and refractive index due to temperature changes (1-way only)	$\sim 10^{-13}$
Coriolis acceleration of the fibre due to velocity of the Earth tides	8×10^{-20}
Euler acceleration of the fibre due to angular acceleration of the Earth rotation	4×10^{-20}
Acceleration of the Earth tides	3×10^{-20}

Table : The corrections depend on processes in the whole fibre and they are calculated per 1000 km of the fibre length [Gersl et al., 2015]

Fibre propagation

The PTB-SYRTE link: Earth rotation signal



It corresponds to a variation of around 5 fs over 12 hours in the time transfer

Fibre propagation

Parameters uncertainty

Parameter	Uncertainty
Time derivative of the fibre temperature (change of length and refractive index; 1-way only)	3×10^{-11} K/s
Fibre velocity in co-rotating frame	0.6 mm/s
Fibre acceleration in co-rotating frame	9×10^{-8} m/s ²
Fibre position	> Earth radius
Earth angular velocity	> 100 % (relative)
Earth angular acceleration	> 100 % (relative)

Table : Input parameters and their maximal uncertainties sufficient for 10^{-18} relative uncertainty in frequency transfer. The values were obtained for situations where the sensitivity of a correction to a parameter is maximized and they are calculated for 1000 km long fibre [Gersl et al., 2015].

Literature I



Gersl, J., Delva, P., and Wolf, P. (2015).

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Soffel, M., Klioner, S. A., Petit, G., Wolf, P., Kopeikin, S. M., Bretagnon, P., Brumberg, V. A., Capitaine, N., Damour, T., Fukushima, T., Guinot, B., Huang, T.-Y., Lindegren, L., Ma, C., Nordtvedt, K., Ries, J. C., Seidelmann, P. K., Vokrouhlický, D., Will, C. M., and Xu, C. (2003).

The IAU 2000 resolutions for astrometry, celestial mechanics, and metrology in the relativistic framework: Explanatory supplement.

The Astronomical Journal, 126(6):2687.



Wolf, P. and Petit, G. (1995).

Relativistic theory for clock syntonization and the realization of geocentric coordinate times.

Astronomy and Astrophysics, 304:653.