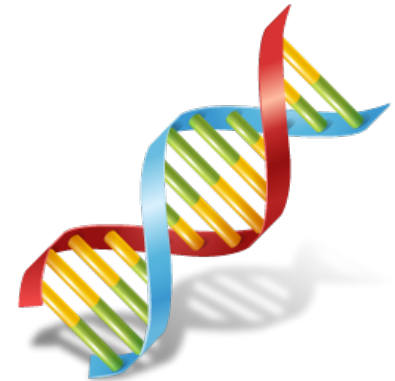
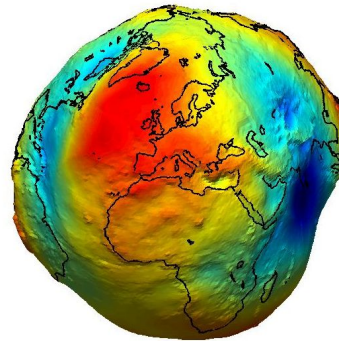
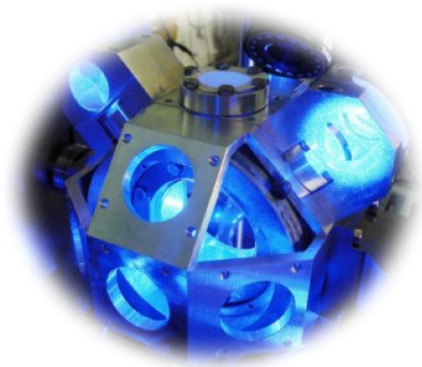


Determination of the geopotential at high spatial resolution with optical clocks



Guillaume Lion

I. Panet, P. Delva, P. Wolf, S. Bize, D. Coulot, C. Guerlin

Institut National de l'Information Géographique et Forestière, LAREG, Université Paris Diderot, Paris, France
SYRTE, Observatoire de Paris, PSL, CNRS, UPMC, LNE, Paris, France

- I. Motivations**
- II. Contribution of atomic clocks for the geopotential determination**
- III. Optimization of the atomic clocks location by genetic algorithms**
- IV. Conclusions and future prospects**

Problems in geodesy...

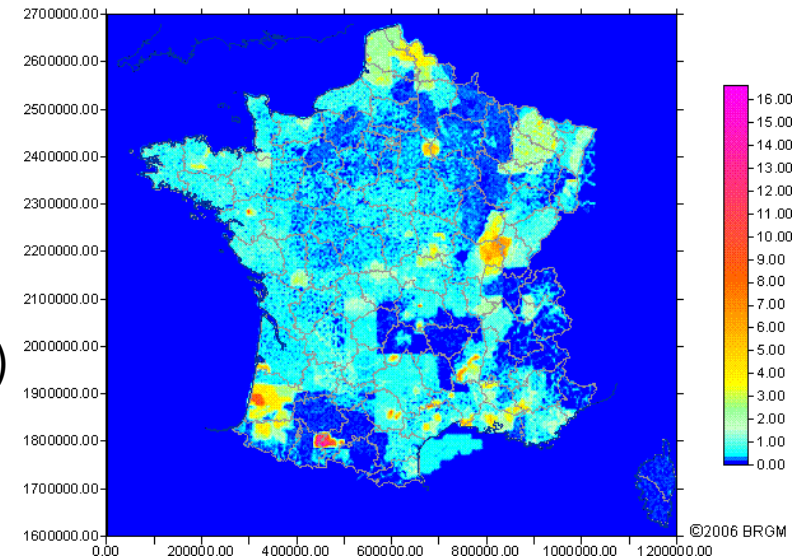
Gravity field and geopotential models

- Quality and heterogeneous coverage of data
 - Quality of the gravity field
→ 0.5 mGal (plains) - 3 mGal (mountains)
 - Quality of the geoid
→ 2 cm (plains) - less than 10 cm (mountains)
- Computation methodology (many parameters)

Definition of vertical frames

- Many and different height systems
- Discrepancy between techniques
- Way to get geopotential heights (by computation...)

Distribution of gravimetric data in France
(Number of stations/km²)



New perspectives to overcome these limitations ?

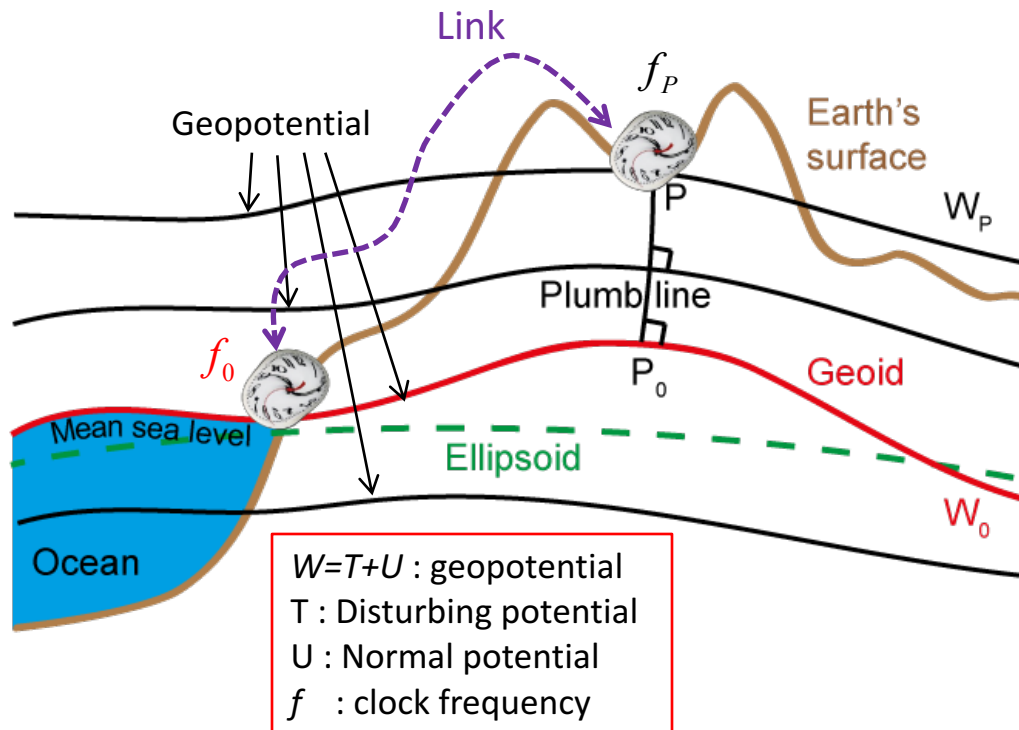
- Geodesy based on optical atomic clocks
- Nice idea to link geopotential and heights to an atomic reference



- Better models of geopotential and gravity field
- Better vertical reference frames

Chronometric geodesy

- GR says that time flows differs for 2 clocks put into different gravitational potential (dilation of gravitational time)
- Clock comparisons permit to determine directly geopotential differences ΔW at the Earth's surface



Relativistic gravitational redshift:

$$\frac{\Delta W}{c^2} = \frac{W_P - W_0}{c^2} \approx \frac{\Delta f}{f}$$

Related to height difference:

$$\Delta W = \frac{\mu_{Earth}}{R_{Earth}} - \frac{\mu_{Earth}}{R_{Earth} + \Delta h} \approx \frac{\mu_{Earth}}{R_{Earth}^2} \Delta h$$

Exactitude $\Delta f / f \approx 10^{-18}$

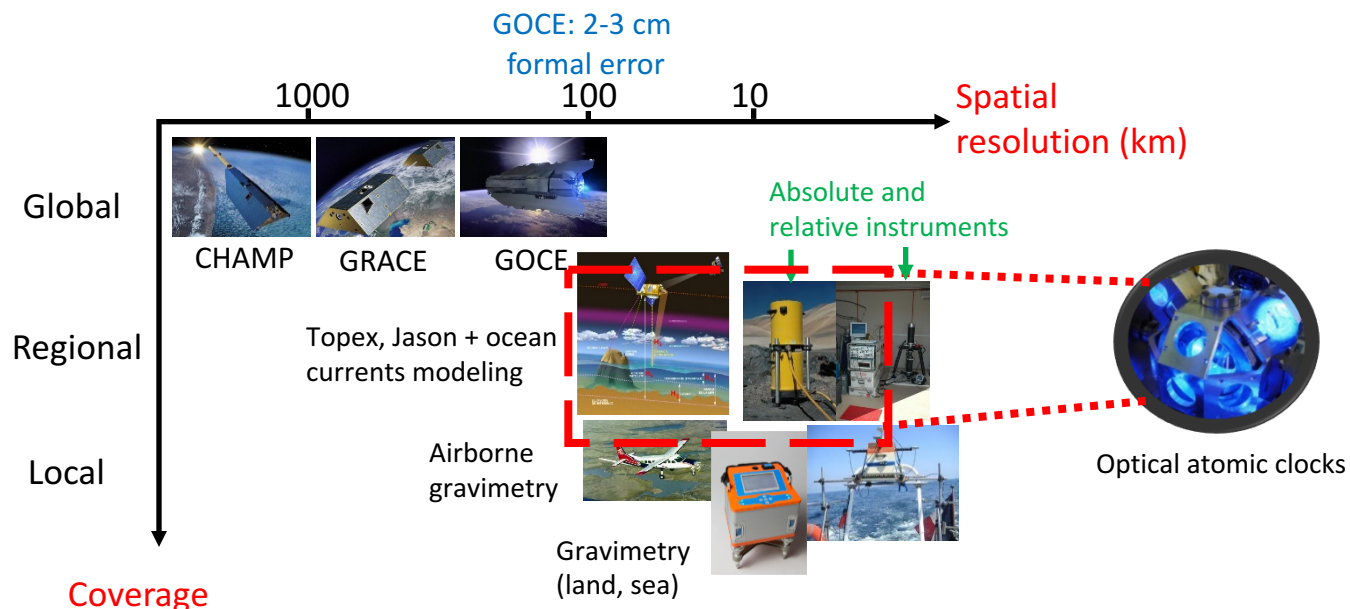


$$\Delta h \approx 1 \text{ cm}$$

$$\Delta W \approx 0.1 \text{ m}^2 \text{s}^{-2}$$

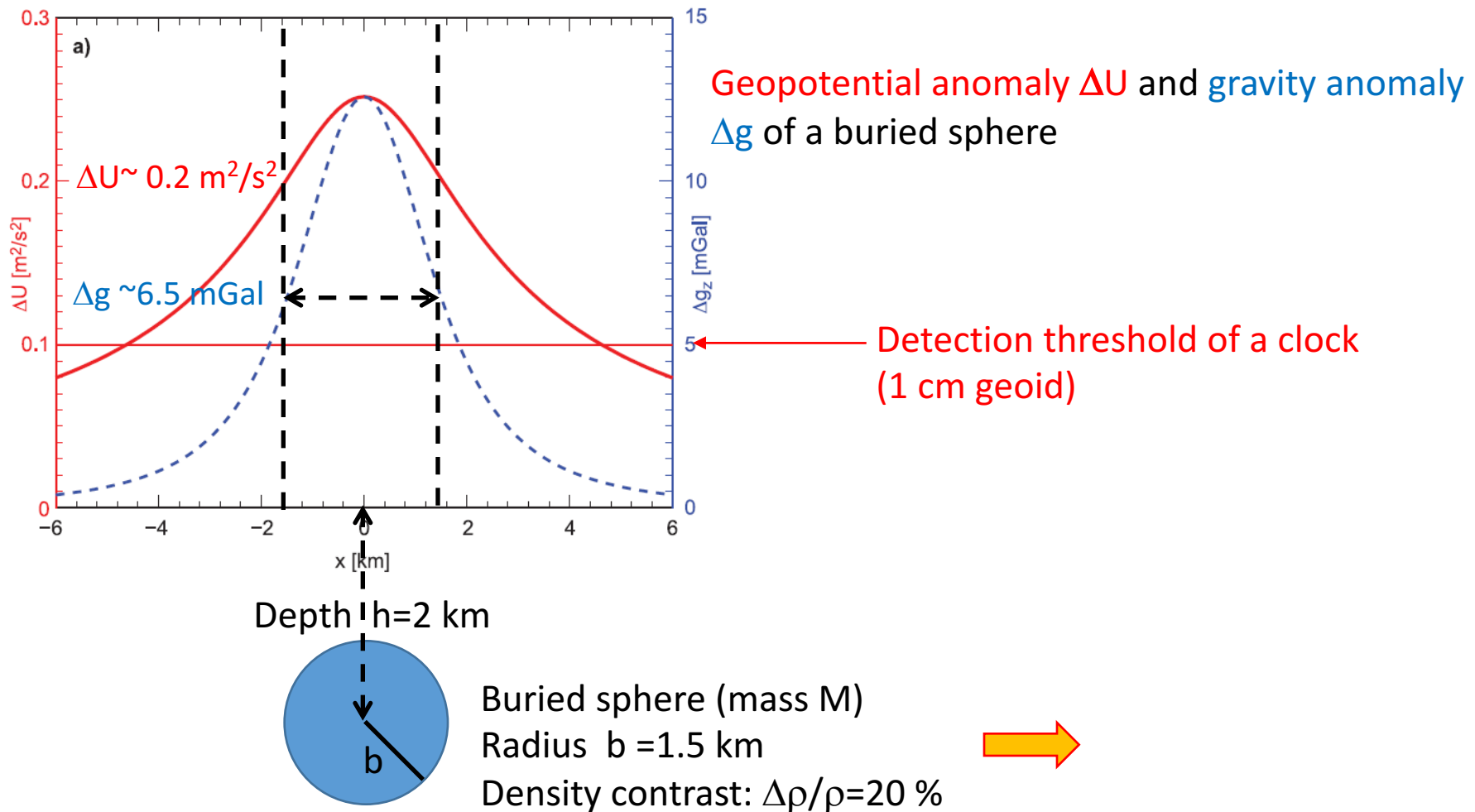
Chronometric observables in geodesy

- **New type of geodetic observable** → Gravitational potential differences
- **Complementary** to gravity and gravity gradients
- **Sensitive to mass variations**, including at depth
- **Better spatial resolution** than satellite techniques
- **Reduction of heterogeneities** in coverage of ground measurements
- **Accuracy closer and closer** to those of other techniques



Geophysical application: gravity and geopotential signal

→ Bondarescu et al. (2012) *Geophysical applicability of atomic clocks: direct continental geoid mapping, JGI*



Aims of the project

➡ Evaluate the contribution of atomic clocks for the determination of the geopotential at high spatial resolution (10 km of resolution)

Study published → Lion G., Panet I., Delva P., Wolf P., Guerlin C., & Bize S. (2017): *Determination of a high spatial resolution geopotential model using atomic clock comparisons*. Journal of Geodesy

➡ How to choose a good coverage of clock measurement points?

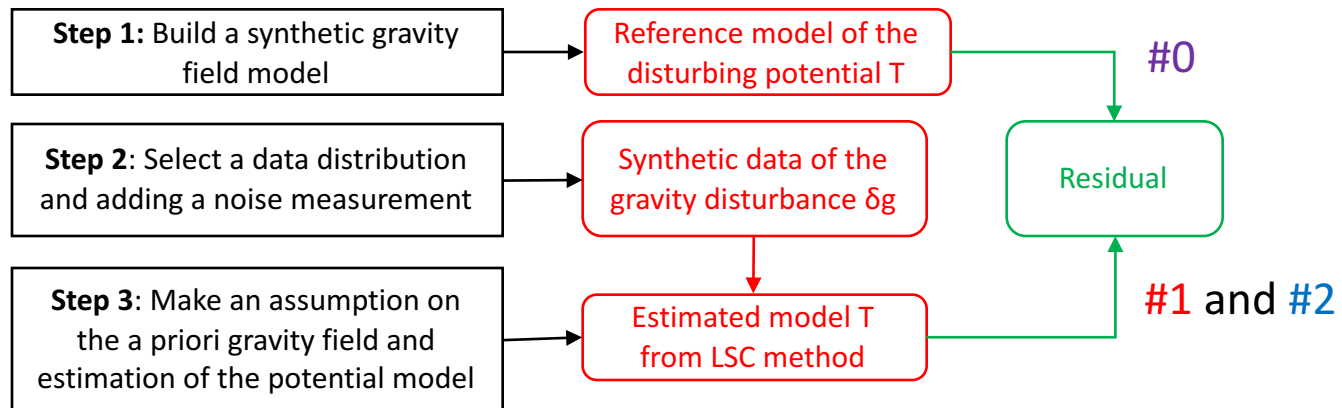
Assumptions


- The work is based on an existing gravity network
- All the geodetic data are synthetics, only their distribution is realistic
- The gravity field is static and isotropic
- We work directly with the disturbing potential $T = W - U$
- Errors on the vertical position is not considered

How to estimate the contribution of clocks data from an gravimetric network ?

Methodology

Tools: Generation, analysis and estimation of a gravity field model



 Evaluation of the contribution of clock measurements by comparing the solutions **#1** and **#2** wrt a reference solution **#0**

- **#1**: only from gravity data
- **#2**: from gravity and potential data

 **T is estimated** on a regular grid interval of 10 km

- Estimation of the potential T on a regular grid using the **Least Squares Collocation** method (LSC) [Moritz, 1980]

Diagram illustrating the Least Squares Collocation (LSC) method for estimating the potential T at point P :

$$\hat{T}_P = C_{T_P, l}^\dagger \cdot (C_{l, l} + C_{\hat{U}\hat{U}})^{-1} \cdot \hat{l}$$

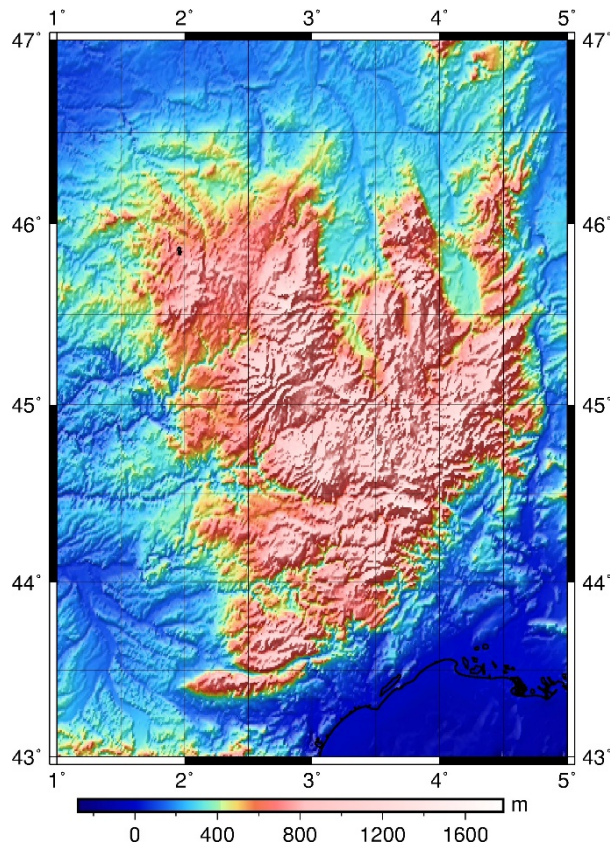
Annotations:

- Estimation of T at point P** (green arrow pointing to \hat{T}_P)
- Covariance matrix between the estimated signal and the signal** (blue arrow pointing to $C_{T_P, l}^\dagger$)
- Covariance matrix of the signal** (black arrow pointing to $C_{l, l}$)
- Covariance matrix of the noise** (black arrow pointing to $C_{\hat{U}\hat{U}}$)
- Signal (data)** (red arrow pointing to \hat{l})

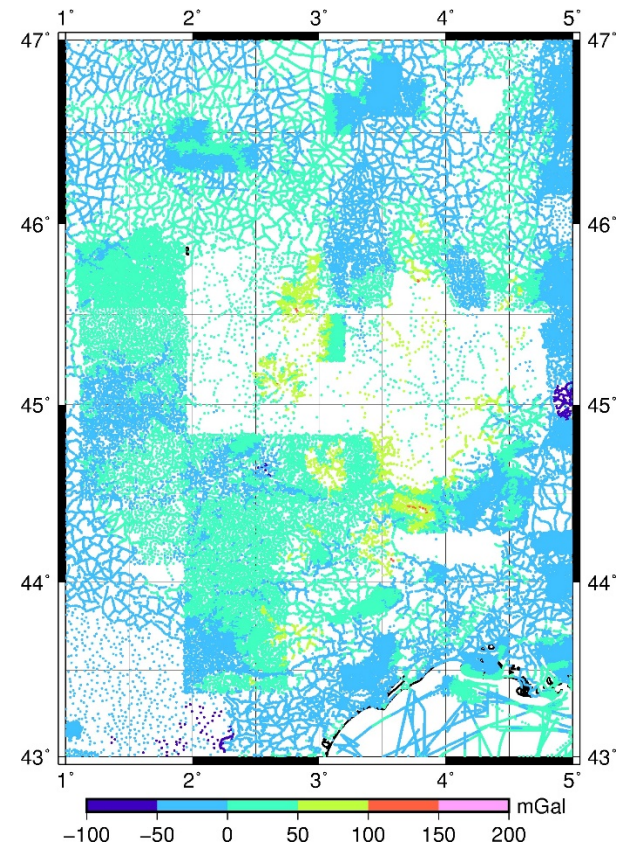
- Assumption on the gravity field a priori regularity
→ Logarithmic 3D covariance function [Forsberg, 1987]
- Inverse covariance matrix is computed using a Cholesky decomposition

Massif central

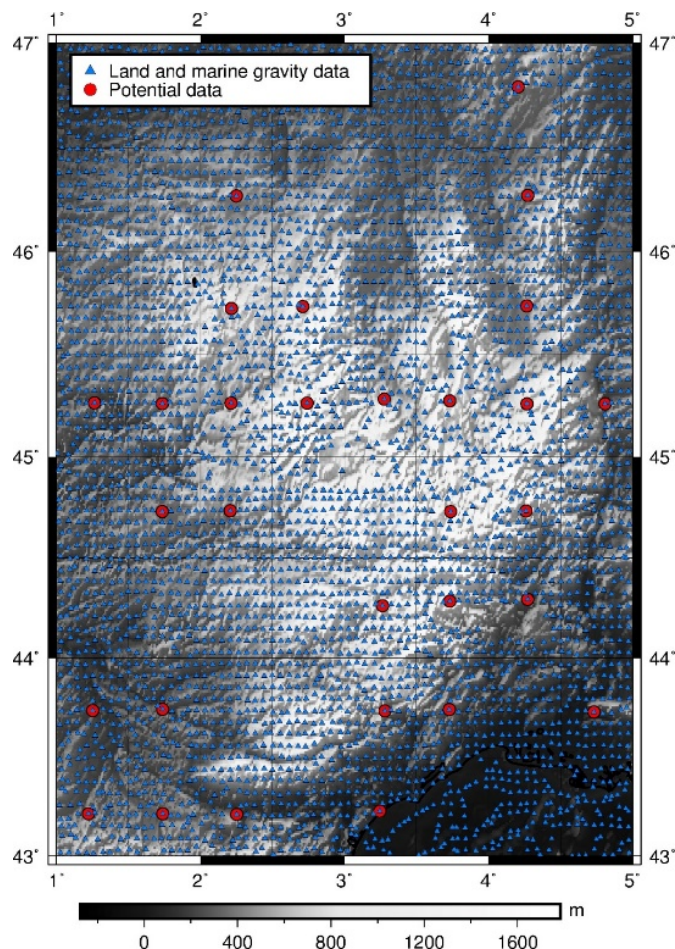
- Moderately mountainous terrain
- Intermediate gravity data coverage: 149522 data (BGI)



Topography



Terrestrial and marine free-air
gravity anomalies



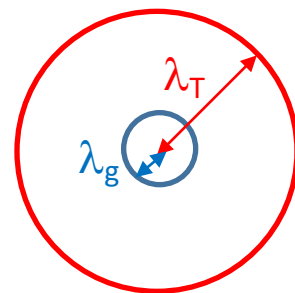
- 4374 reduced gravity data δg
→ noise = 1mGal
- 33 potential data T
→ noise = 0.1 m²/s²

How to select the gravi points ?

- Data reduction from the ~150000 locations
- Distance between each point ~6.5 km
- Each point is weighted (number of real points in the vicinity)

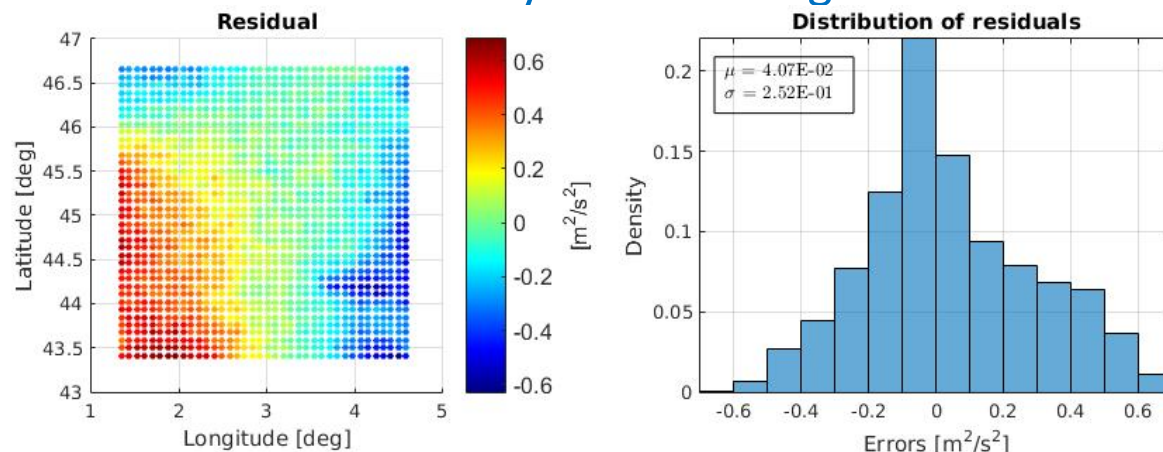
How to select the clock points ?

- T more sensitive to medium wavelengths λ than δg
- The location of the clock points is chosen to better complete the gravity network
- T at same location as δg
- Red points are an example of “handmade coverage” (not optimized)



White noise is added to the perfect synthetic data

Reconstruction of T only from data δg



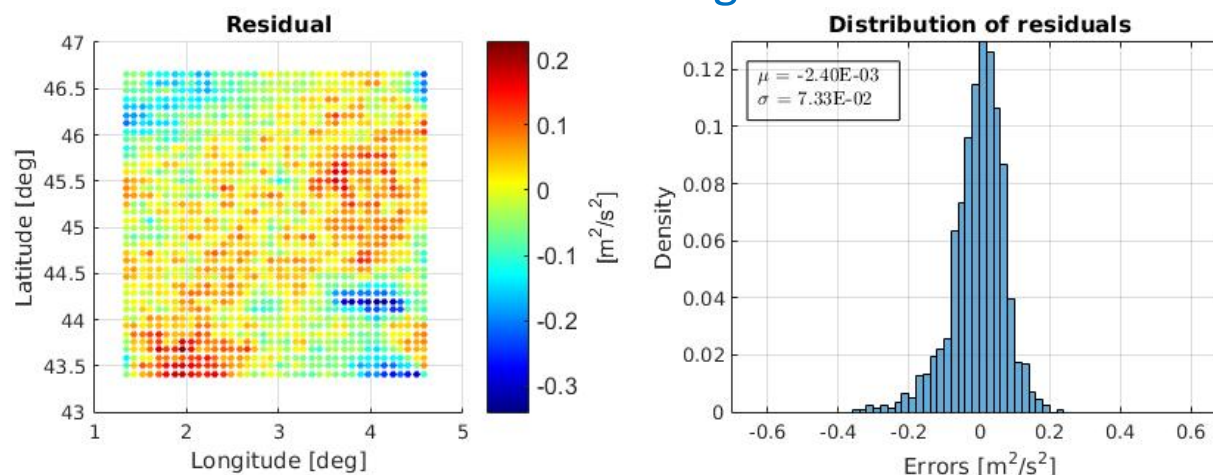
Bias: $-0.04 \text{ m}^2/\text{s}^2$

→ 4mm

RMS: $0.3 \text{ m}^2/\text{s}^2$

→ 3cm

Reconstruction of T from data $\delta g + T$



Bias: $-0.002 \text{ m}^2/\text{s}^2$

→ 0.2mm

RMS: $0.1 \text{ m}^2/\text{s}^2$

→ 1cm

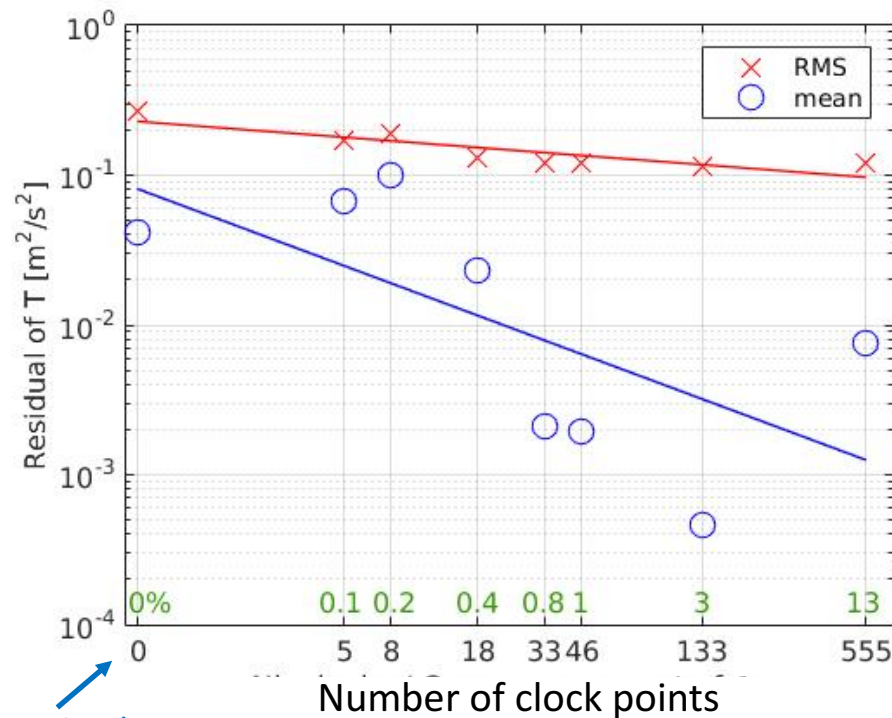


- Reduction of the bias
- Tighter dispersion

Results from *Lion et al. (2017)*

Effect of the clock coverage

Results from *Lion et al. (2017)*



← 1 cm

Around 30 clock points permit to reduce the bias by 1 order of magnitude

← $\frac{\text{Nb clock points}}{\text{Nb gravi points}} \times 100$

Only gravity data

Effect of the noise levels (m^2/s^2)

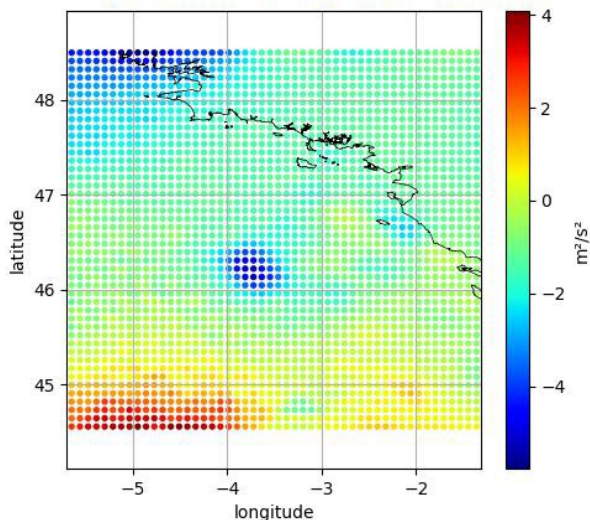
$\sigma_T \backslash \sigma_{\delta g}$	1 mGal		0.1 mGal		0.01 mGal	
	μ	σ	μ	σ	μ	σ
No clock	2.2×10^{-1}	3.9×10^{-1}	2.1×10^{-1}	4.2×10^{-1}	2.1×10^{-1}	4.2×10^{-1}
$1 \text{ m}^2 \text{ s}^{-2}$	1.4×10^{-1}	3.4×10^{-1}	1.2×10^{-1}	3.3×10^{-1}	1.2×10^{-1}	3.3×10^{-1}
$0.1 \text{ m}^2 \text{ s}^{-2}$	6.8×10^{-2}	1.7×10^{-1}	4.7×10^{-2}	1.5×10^{-1}	1.7×10^{-2}	1.6×10^{-1}

🌐 Test of our method on the littoral zone (Atlantic coast)

→ Hugo Lecomte (master internship)

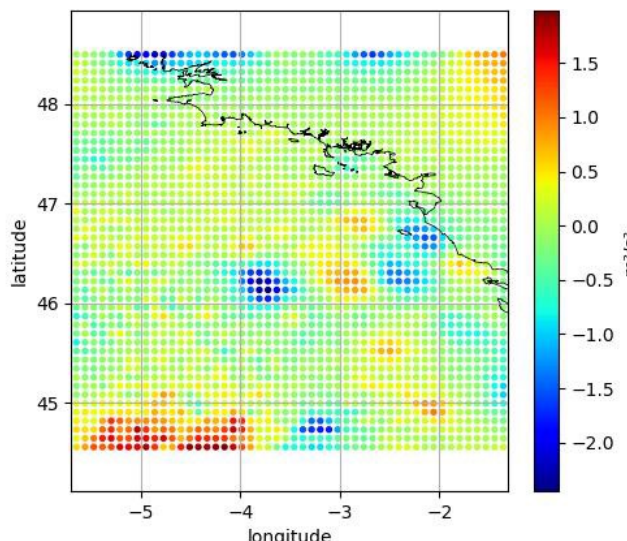
Mono-covariance
Residual of T ...

only from data δg



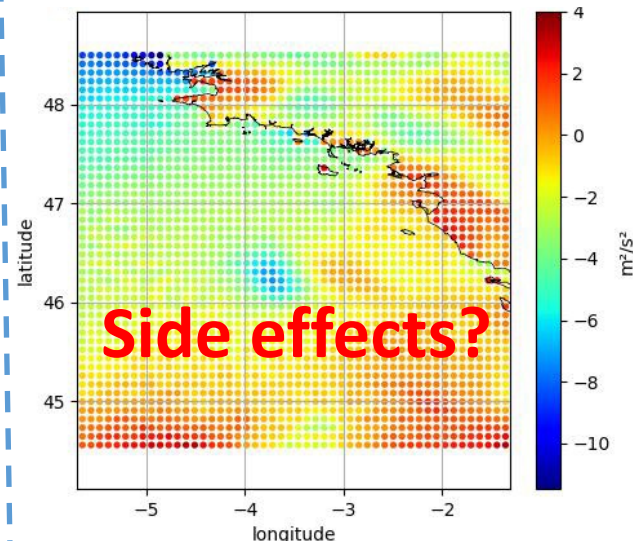
Bias: $-0.9 \text{ m}^2/\text{s}^2$
RMS: $1.3 \text{ m}^2/\text{s}^2$

from data $\delta g + T$

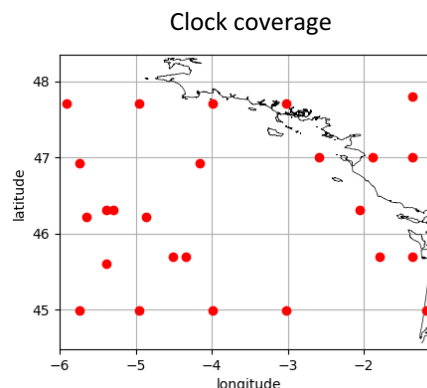


Bias: $-0.06 \text{ m}^2/\text{s}^2$
RMS: $0.5 \text{ m}^2/\text{s}^2$

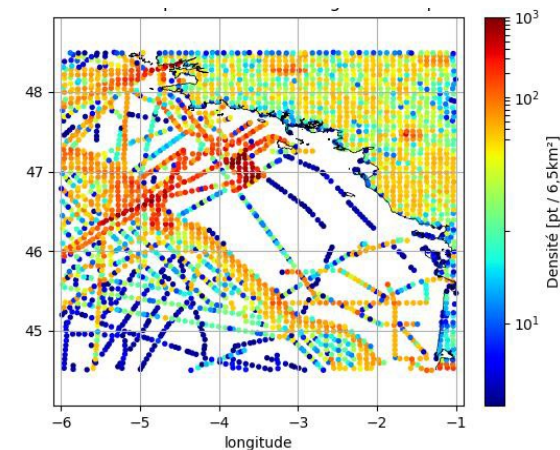
from data δg



Bias: $-1.1 \text{ m}^2/\text{s}^2$
RMS: $1.6 \text{ m}^2/\text{s}^2$



Terrestrial and marine free-air gravity anomalies



Multi-covariance
Residual of T ...



**How to choose location of the clock points
to improve the reconstruction of T ?**

Principle of Genetic Algorithms (GA)

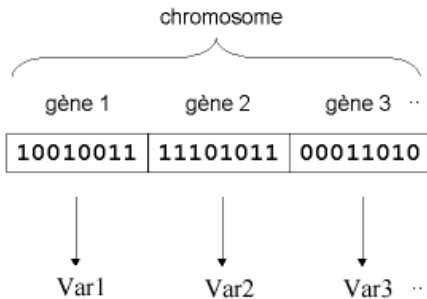
- Based on Darwin's theory of **natural selection** and **modern genetics**:
 - **Natural selection**: best adapted individuals to their environment ("fit") survive longer and reproduce more easily
 - **Genetic**: the laws of heredity use notions of crossover-mutation. The evolution and the adaptation of the species is explained by referring the lucky factor :
 - ⇒ species evolve without a predetermined goal
 - ⇒ random spontaneous births



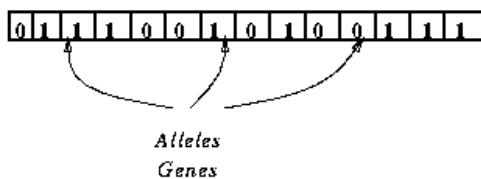
- **Idea**: solving complex optimization problems by simulating the process of biological evolution

🧬 Elements of a Genetic Algorithms (GA)

COMPUTER

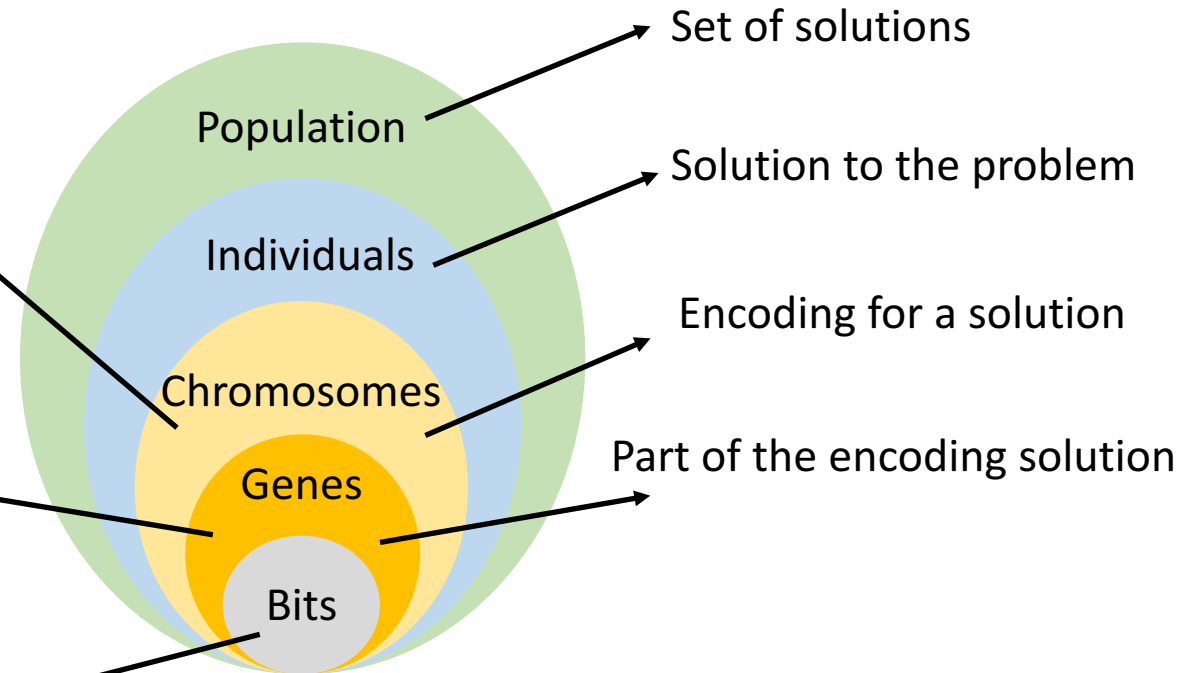


Position of 1s and 0s control the characteristic of the problem



Binary encoding (0,1)

NATURE



- Encoding depends entirely on the problem heavily !
- Binary is the most common method

In practice

List of location points and data:

LAT (deg)	LON (deg)	H (m)	T (m ² /s ²)	Noise T	ID	Binary string
45.621496	4.040226	535.515300	-0.145871	0.100000	53	00110101
46.111996	2.898492	524.439286	0.162066	0.100000	149	10010101
...					...	
44.737321	1.439303	328.054267	1.241396	0.100000	581	11001001

- Each data point is recognized by an integer ID (line number for example)
- The ID of the gene is encoded in binary string
- Decoding function d for a gene of length l is:

$$x = d(A) = \sum_{i=1}^l a_i 2^{l-i-1}$$

Ex: $A = \{a_i\} = \{1, 0, 1, 1\} \rightarrow x = 11$

- Accuracy of the decoding s (depends on the number of points to be tested):

$$X' = x_{min} + x \frac{x_{max} - x_{min}}{2^s - 1}$$

Ex: if we have 1056 potential locations

→ ID = $[x_{min} = 0; x_{max} = 1055]$

→ Encoding of the IDs with $s=11 \Rightarrow 2^{11} = 2048$ bits

→ The GA pulls the gene $A = \{1111010\} \rightarrow x = 122 \rightarrow X' = 62.8 \Rightarrow 62$

→ Accuracy of 1/2048 in the search space

Classical scheme associated with GAs

0101011
1101110
...
0100011

Initial random population

Generation 0

Evaluation

Parents

Cost function
- Criteria
satisfied?

YES

Best individuals

Solutions

1001001
0101100

NO

Randomly pick 2
parents

reproduction

Crossover

Parents

0101011
1101110

1 point
crossover

0101110
1101011

used

2 points
crossover

0101111
1101010

Offspring

Mutation

0101011

0111011

Next
generation



Pareto dominance

- Context

Minimization of p objective functions f_i defined on $\Omega \in R^n$

- Dominance in the sense of Pareto

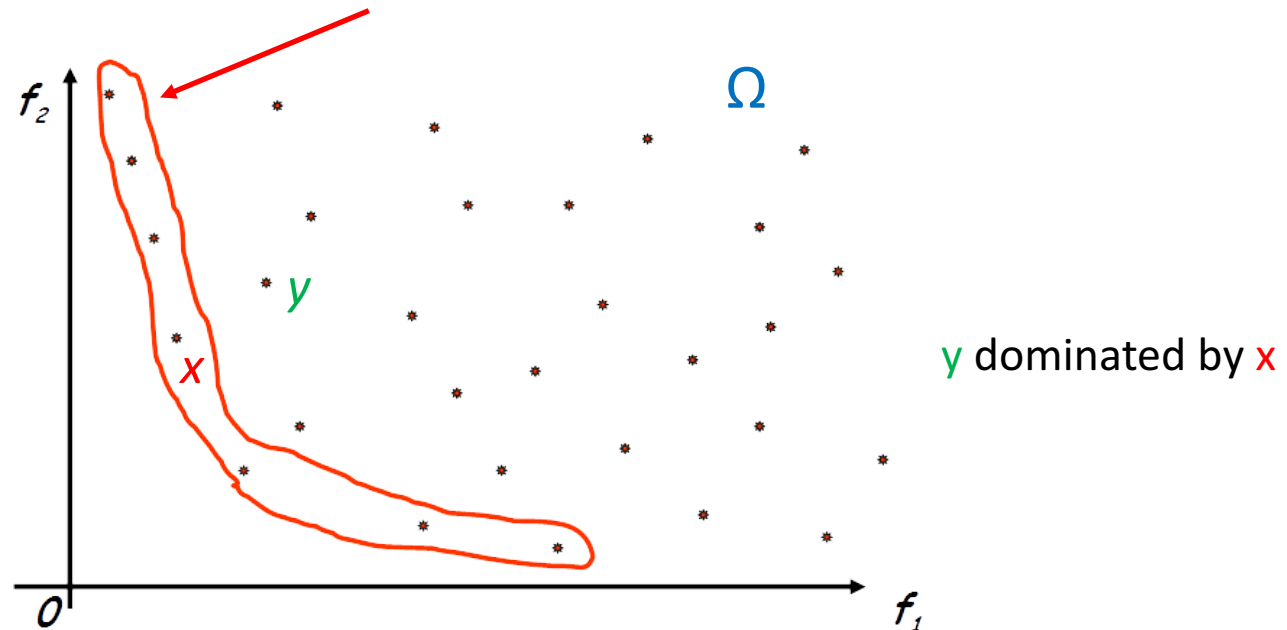
Partial order relationship defined by

$$x < y \text{ ssi } \begin{cases} \forall i \in [1, p] f_i(x) \leq f_i(y) \\ \exists j \in [1, p] f_j(x) < f_j(y) \end{cases}$$

- Pareto front-set

Set of non-dominated solutions of Ω

Image in the search space = **Pareto front**



ε-MOEA

- ε-MOEA : ε Multi-Objective Evolutionary Algorithm
- Developed by **Pr. Deb (2003)**
- Recent genetic algorithm, faster and with very good convergence compared to other schemes (McMc, NSGA-II, etc.)
- **Characterized by 2 main principles:** ε-dominance and the archive

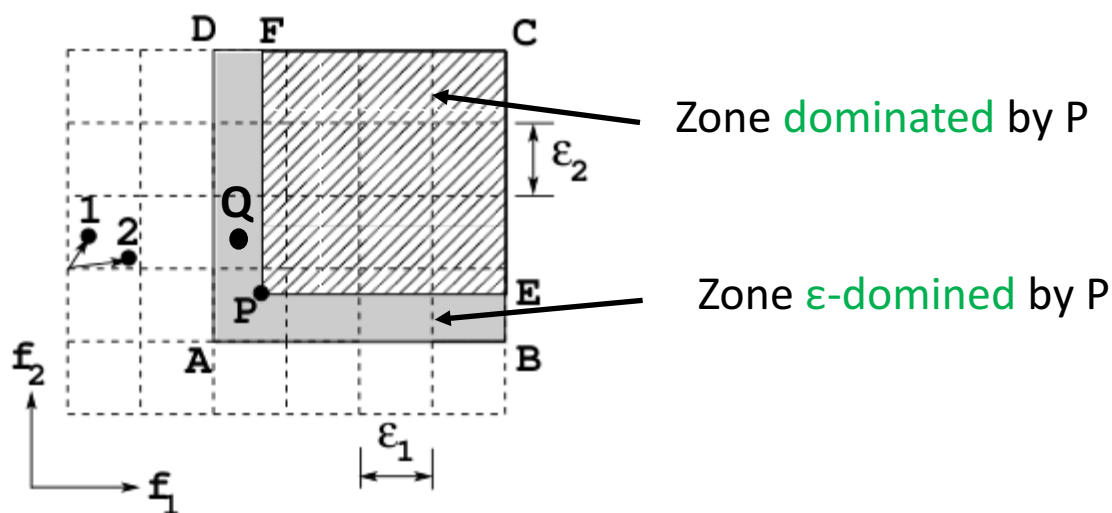
ε-dominance [Laumanns et al., 2003]

- Modified dominance relation (ϵ_i , real p fixed)
- Ex: for 2 objectives

$$x <_y \text{ssi} \begin{cases} \forall i \in [1, p] f_i(x) \leq f_i(y) + \epsilon_i \\ \exists j \in [1, p] f_j(x) < f_j(y) + \epsilon_j \end{cases}$$

ε = tolerance on the value of the objectives (resolution):

- allows diversity and influences the size of the Pareto front
- eliminates points that are too close together

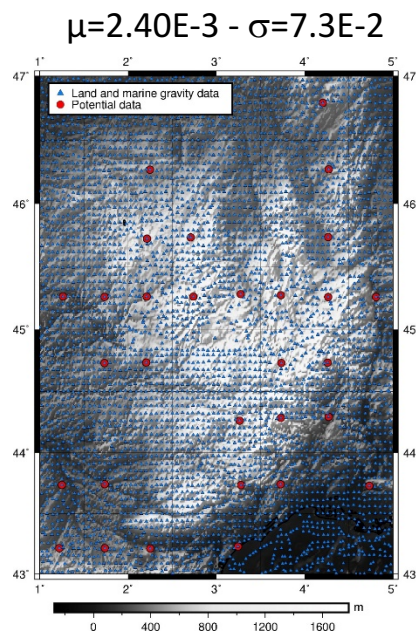


Although solutions P and Q are non-dominated with each other,
P ε-dominates Q since the latter belongs to the region ε-dominated by P

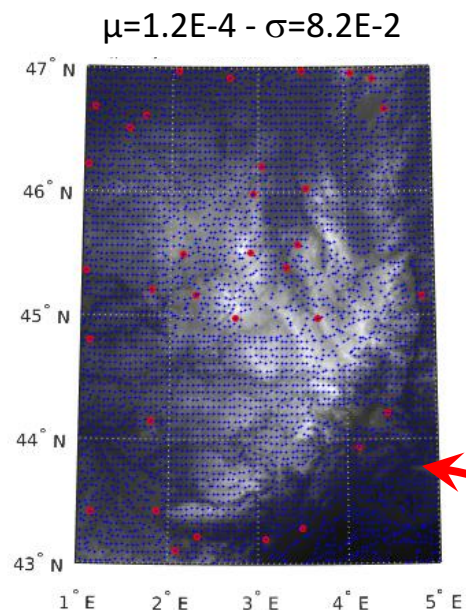
BINARY CODING + chromosomes with fixed lengths

🧬 Optimization of the position of 33 clocks

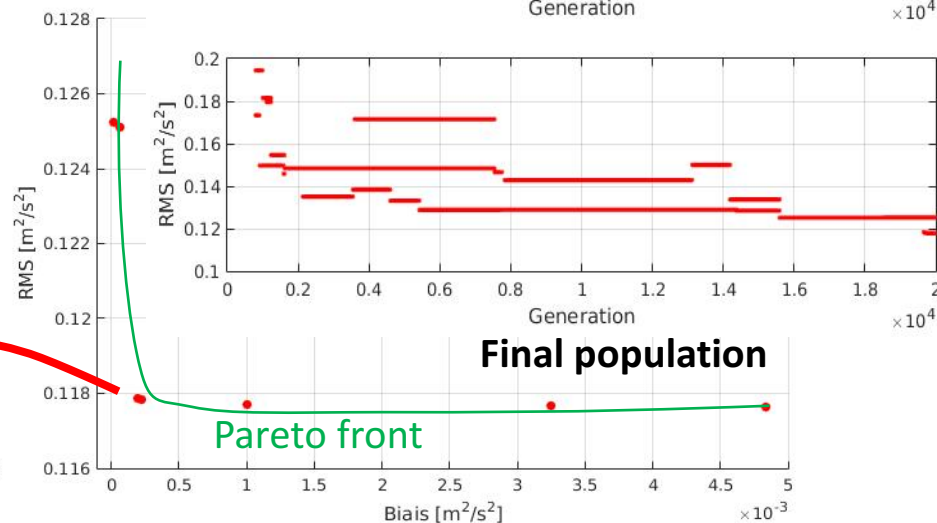
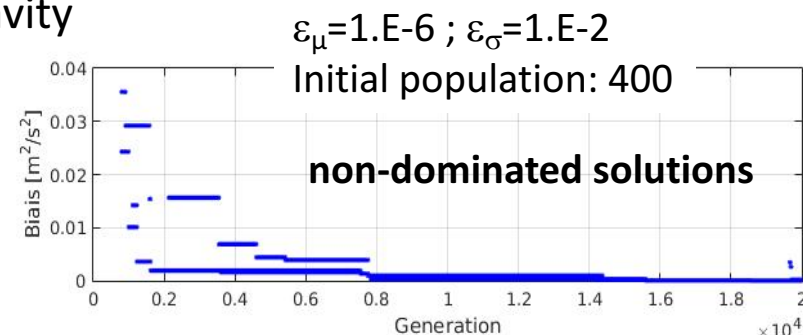
- **Cost function (objectives)**
 - Minimization of the reconstruction residuals on T (bias $|\mu|$ and STD σ)
- **Constrains**
 - Clock point at the same place as a gravity point
 - Clock point in an area poorly covered by gravity
 - Minimum distance between 2 clock points
 - Clock points on ground



“Handmade”



“Genetically” modified



Integer encoding with variable length chromosomes

- Need to **customize the GA to suit the problem** or class of problems under consideration
- Integer encoding** → to increase the efficiency of the genetic algorithm, i.e. better convergence: each clock point is associated with its position in the sequence
- Variable length strings** → to minimize the number of observables
- Set a limit on maximum length of chromosome (e.g. 16 location points)

Two-point crossover

Parent A

1	5	6	8	10	15	16	23	26	0	0	0	0	0	0	0
---	---	---	---	----	----	----	----	----	---	---	---	---	---	---	---

Parent B

3	7	8	9	11	13	14	19	21	22	23	27	31	0	0	0
---	---	---	---	----	----	----	----	----	----	----	----	----	---	---	---

Offspring

1	5	6	8	10	13	14	19	21	22	0	0	0	0	0	0
---	---	---	---	----	----	----	----	----	----	---	---	---	---	---	---

3	7	8	9	11	15	16	23	26	0	23	27	31	0	0	0
---	---	---	---	----	----	----	----	----	---	----	----	----	---	---	---

String completed by 0s

Genes are then sorted in ascending order

Mutation

3	7	8	9	11	15	16	23	26	27	31	0	0	0	0	0
---	---	---	---	----	----	----	----	----	----	----	---	---	---	---	---

3	7	28	9	11	15	16	23	0	27	31	0	0	1	0	0
---	---	----	---	----	----	----	----	---	----	----	---	---	---	---	---

change

annihilation

creation

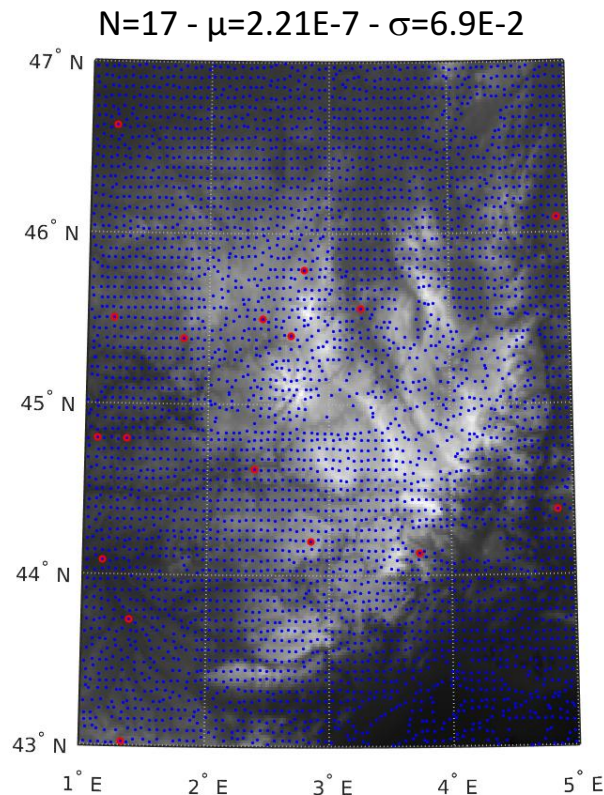
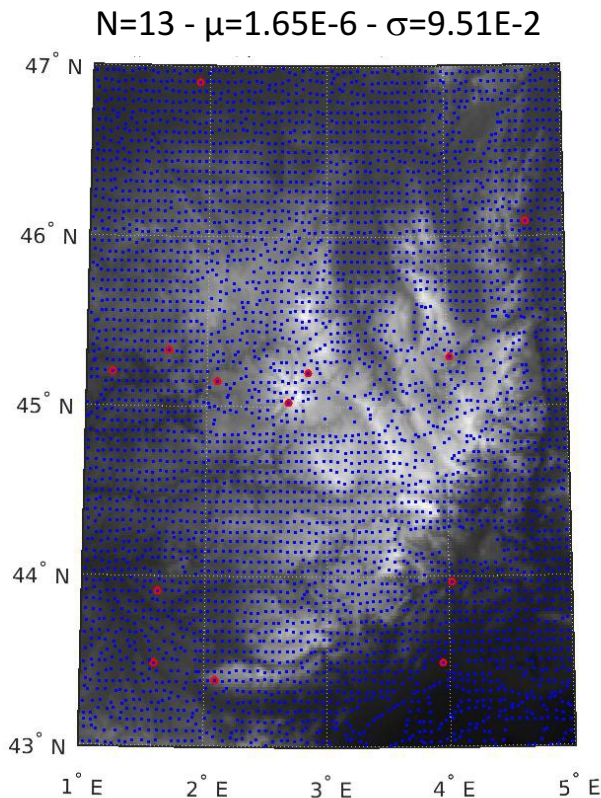
INTEGER CODING + chromosomes with variable lengths

[Coulot et al. 2015]

- Cost function (objectives)

→ Minimization of the reconstruction residuals of T (bias $|\mu|$ and STD σ)

→ Minimization of the number of clock point location $N = [5; 50]$



$\varepsilon_\mu=1.\text{E-}6$; $\varepsilon_\sigma=1.\text{E-}2$; $\varepsilon_N=0.5$

Initial population: 400

Number of generation : 20000

CPU time: about 8 days

- Small number of clock points
- Reduction of the bias by 4 orders of magnitude than previous scenarios



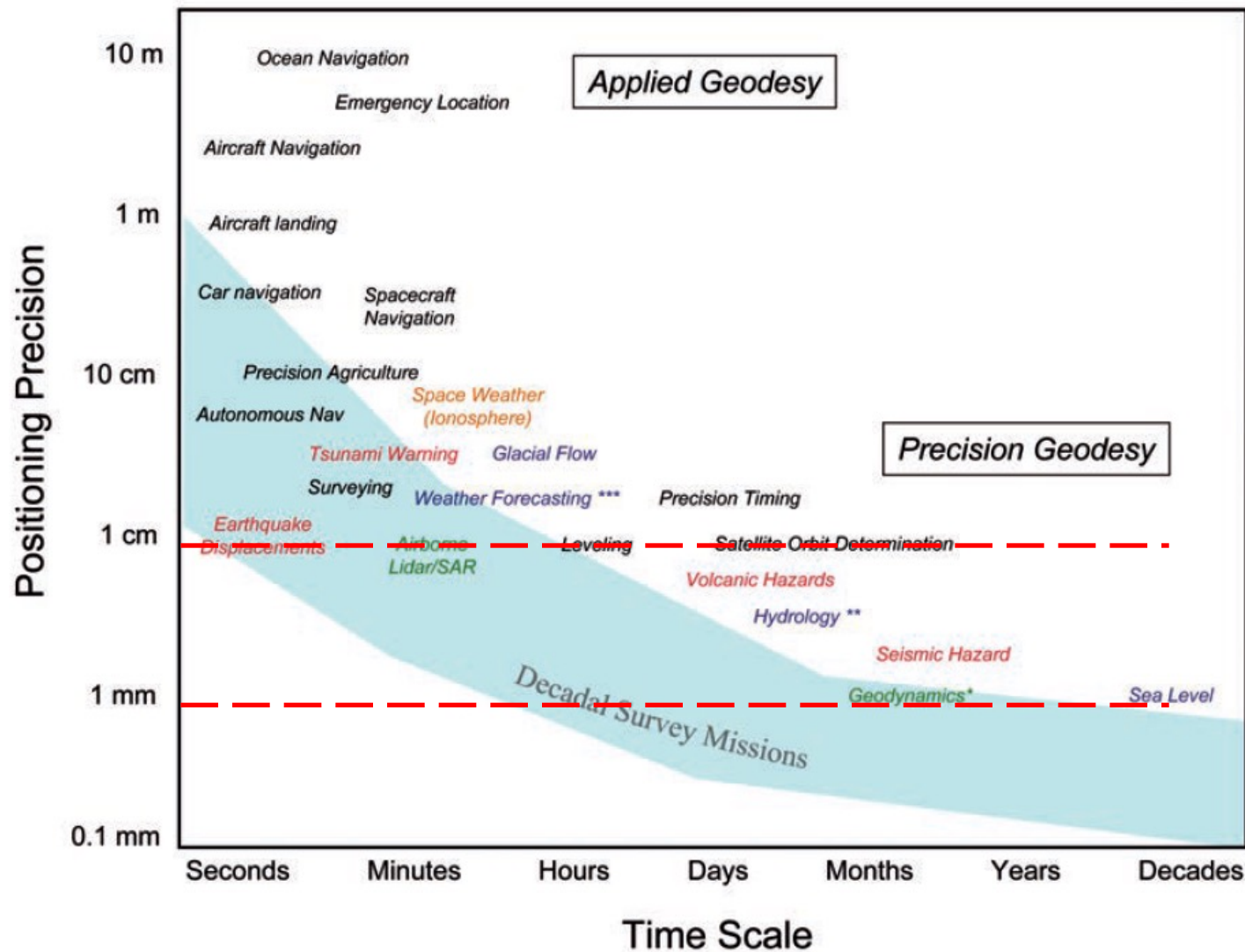
Better coverage

(article in process...)

- ✓ Clocks provide complementary information to surface and satellite data, particularly in areas poorly covered by gravity data
- ✓ In our study area, around 30 clock points improve the potential reconstruction (by more than 2 orders of magnitude) and improve accuracy bias (by a factor 3)
- ✓ Results also confirmed in Alps-Mediterranean region, Bretagne and Atlantic coast
- ✓ A GA is a powerful tool to find optimal data coverages (of any kind): smaller residuals on T with 50% less clock points wrt the first results

- Cost function of the GA evaluated locally in sub-areas (in progress...)
- Use of the formal errors of the LSC method instead of a computed reference solution (in progress...)
- Spatially correlated error patterns (colored noise) to identify biases in dataset
- Test different a priori models (isotrop and non-isotrop)
- Optical clocks on the littoral zone ? On board ship ?
- Sensitivity of clocks to the vertical displacement
- Can optical clocks ($10^{-19} - 10^{-20}$) could detect geodynamic processes inducing Earth deformations ? (mass transport, seismic cycle, surface loading, hydrological cycle, ...)
- **Other ideas ?**

National Report Council (NRC) on Precise Geodetic Infrastructures, 2010



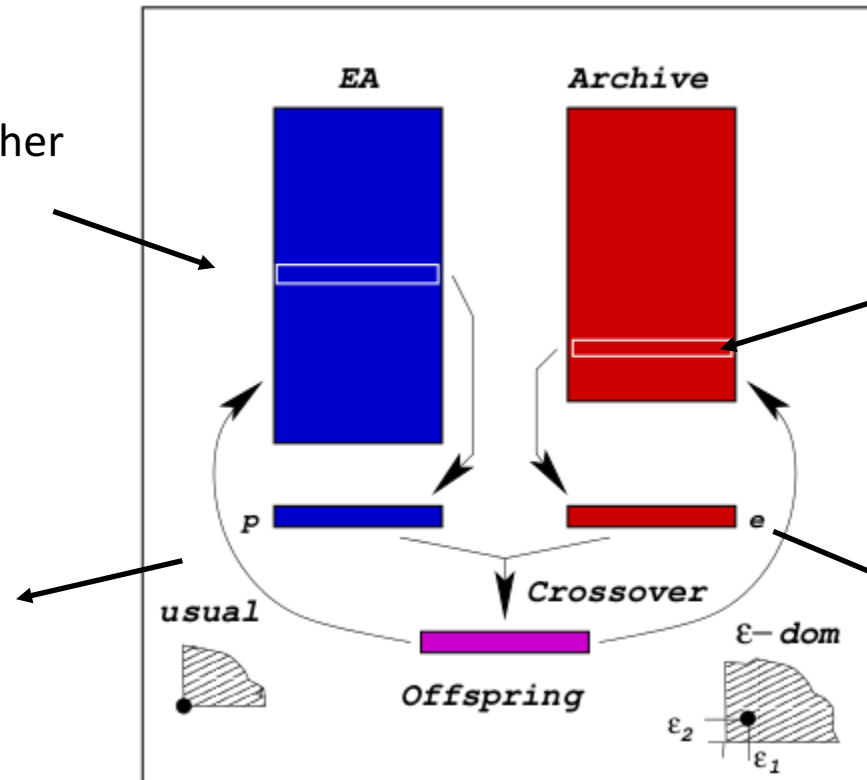
ε-MOEA

2 random individuals

Tournament:

- one dominates the other
- otherwise, chance

Comparison dominance



a: archive
o: offspring
p: parent

1 random individual

Comparison
ε-dominance

Comparison of o with all p 's of P (population size kept constant)

- \underline{o} dominates one of the \underline{p} \rightarrow \underline{o} replaces \underline{p}
- \underline{o} is dominated by a \underline{p} \rightarrow \underline{o} is not retained
- Otherwise, \underline{o} replace a randomly chosen \underline{p}

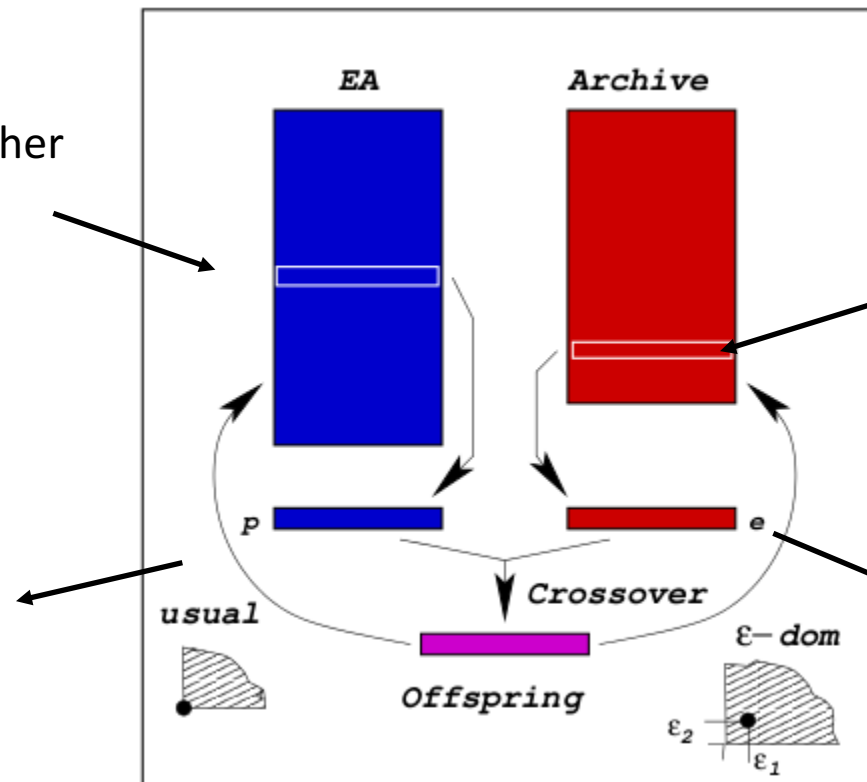
ϵ -MOEA

2 random individuals

Tournament:

- one dominates the other
- otherwise, chance

Comparison dominance



a: archive
o: offspring
p: parent

1 random individual

Comparison
 ϵ -dominance

Comparison with the archive (variable size controlled by the value of ϵ) of \underline{o} with all of A's

- \underline{o} ϵ -dominates one of the \underline{a} \rightarrow \underline{o} replaces \underline{a}
- \underline{o} is ϵ -dominated by \underline{a} \rightarrow \underline{o} is not retained
- Otherwise, 2 cases can occur:
 - \rightarrow \underline{o} is in the same box as \underline{a} of A : we compare \underline{o} and \underline{a} with dominance
 - \rightarrow None of A shares box with \underline{o} \rightarrow \underline{o} is accepted

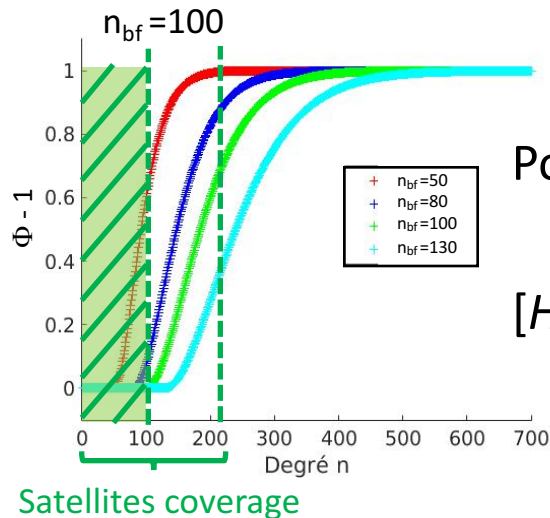
Step 1



Generation of the anomalies δg and T at the Earth's surface (ellipsoidal approximation)

🌐 Global gravity model at 10 km resolution ($n=2000$)
→ EIGEN-6C4 [Förste et al., 2014]

🌐 Removal of low frequencies

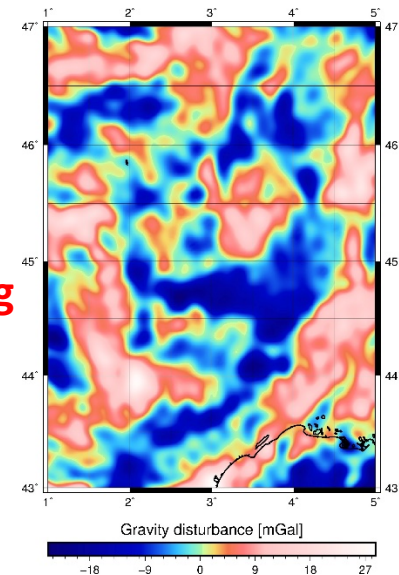


Poisson wavelet spectrum:
 $\Phi(t) \propto t^3 \exp(-t)$
[Holschneider et al., 2003]

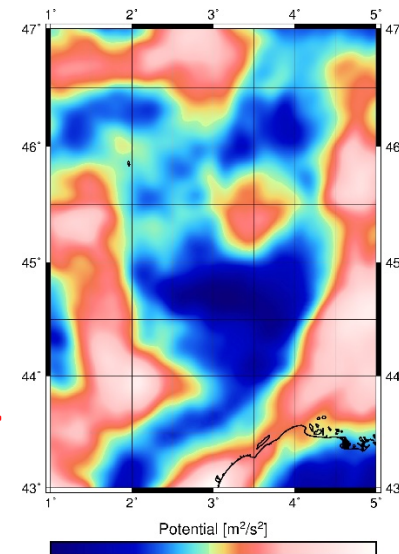
🌐 Subtract the terrain effects
→ dV_ELL_RET2012 [Claessens and Hirt, 2013]

Synthetic field (reference model)

δg



T



🌍 Assumption on the gravity field a priori regularity

→ Logarithmic 3D covariance function [Forsberg, 1987]

$$C'(x, y, z_1 + z_2) = f \sum_{i=0}^3 \alpha_i C(x, y, z_i) \quad \text{with}$$

$$x = x_2 - x_1; \quad y = y_1 - y_1$$

$$z_i = z_1 + z_2 + D_i$$

$$\alpha_i = \{1, -3, 3, 1\}$$

$$D_i = D + iT$$

$$f = C_0 \log^{-1} \left(\frac{D_1^3 D_3}{D_0 D_2^3} \right)$$

- 3 parameters to be estimated:

- C_0 : variance of δg
- D and T : shallow and compensate depth parameter

- Auto-covariance function (ACF) for δg :

$$C = C_{zz} = \frac{\partial}{\partial z_1} \frac{\partial C_{TT}}{\partial z_2} = -\log(D_i + r_i) \quad \text{with} \quad r_i = \sqrt{d^2 + D_i^2}; \quad d^2 = x^2 + y^2$$

- ACF for T:

$$C = C_{TT} = \iint C_{zz} dz_1 dz_2 = \frac{3}{4} z_i r_i + \left(\frac{r_i^2}{4} - \frac{3}{4} z_i^2 \right) \log(z_i + r_i)$$

- Auto-correlation function between T and δg :

$$C = C_{TT_z} = r_i - z \log(z_i + r_i)$$

Step 3

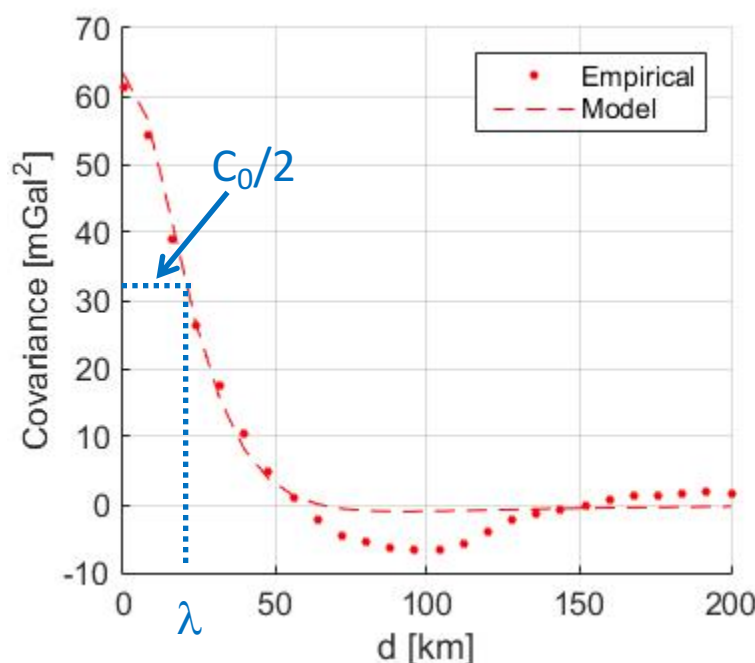
🌐 Assumption on the gravity field a priori regularity

→ Logarithmic 3D covariance function [Forsberg, 1987]

- C_0 : variance of δg
- D and T : shallow and compensate depth parameter

3 parameters to be estimated

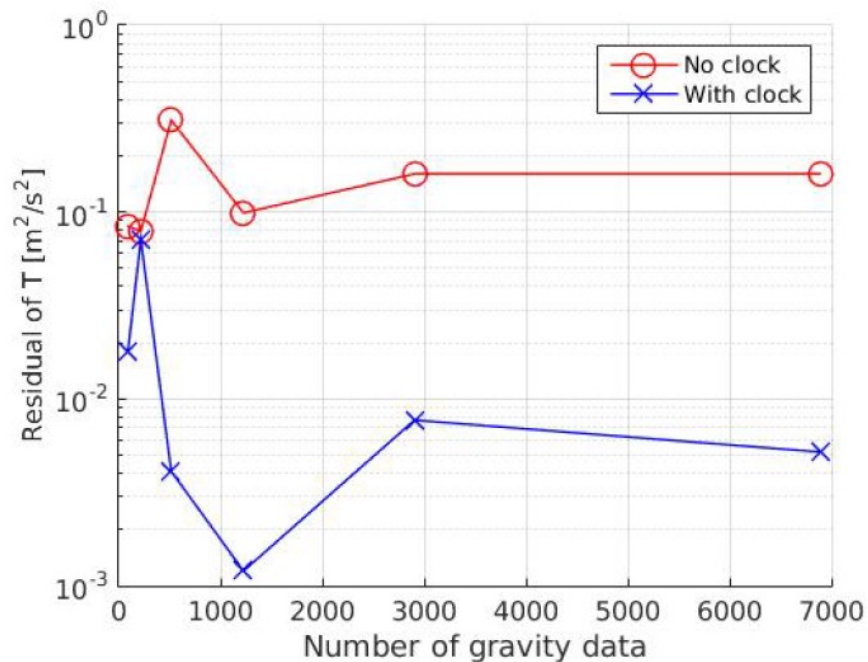
Parameters estimation of the covariance function from the empirical ACF δg



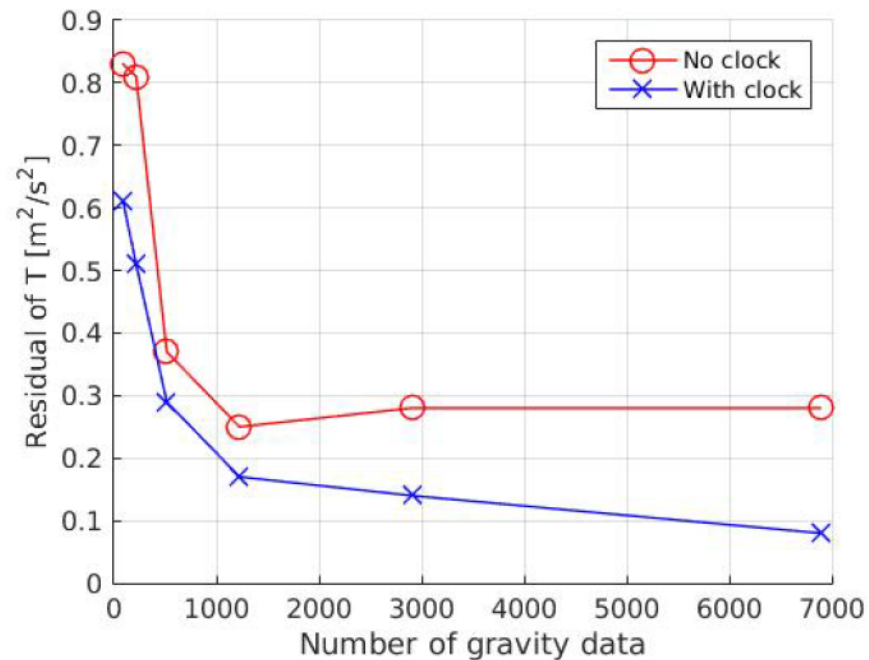
- $C_0 = 63.35 \text{ mGal}^2$ (estimated)
- $D = 24 \text{ km}$
- $T = 15 \text{ km}$

Correlation length: $\lambda \approx 21 \text{ km}$

Effect of the number of gravity data



(a) Absolute value of the mean.



(b) Rms.

→ Adding 38 clocks measurements reduces significantly the trend from the modeling error

→ Increasing the gravity data resolution has more impact on the interpolation errors at medium to smaller scales

→ Adding clocks leads to a further reduction of the rms of residuals